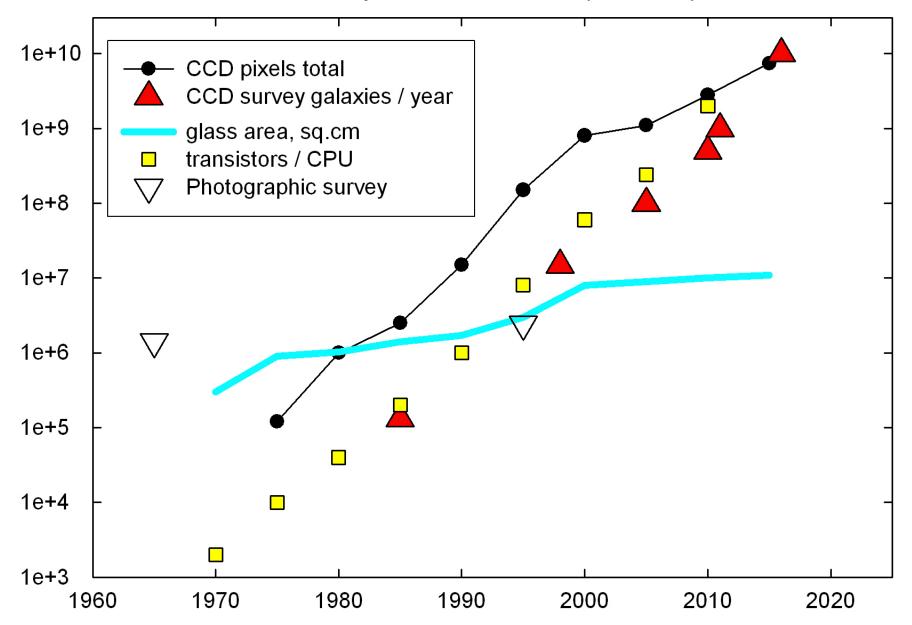
LSST: Petascale opportunities and challenges

Tony Tyson, University of California, Davis

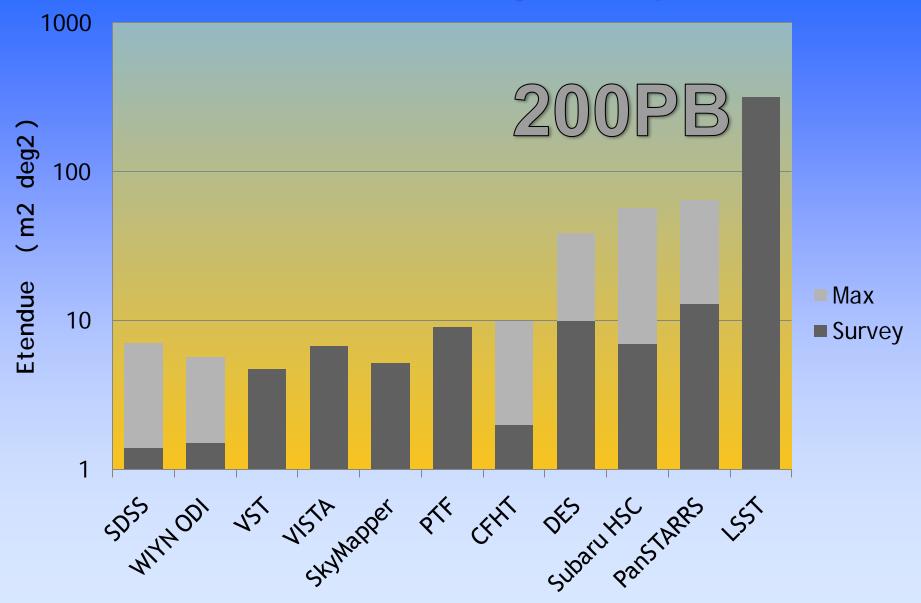


Trends in Optical Astromomy Survey Data





Relative data volume from survey telescopes & cameras



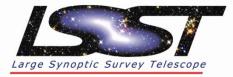
The new sky

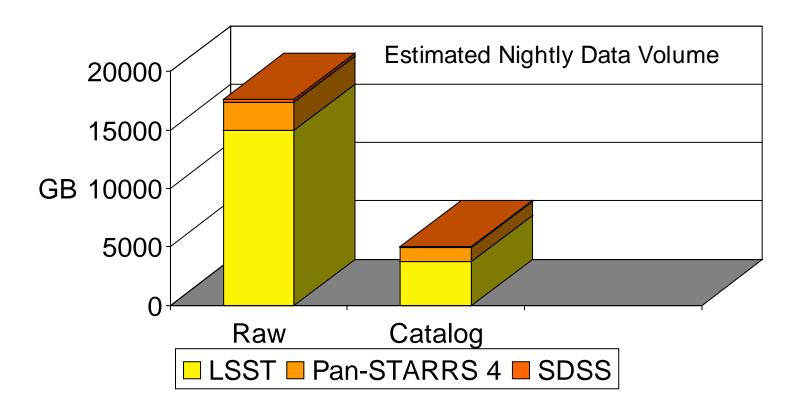


Large Synoptic Survey Telescope



Data volumes & rates are unprecedented in astronomy

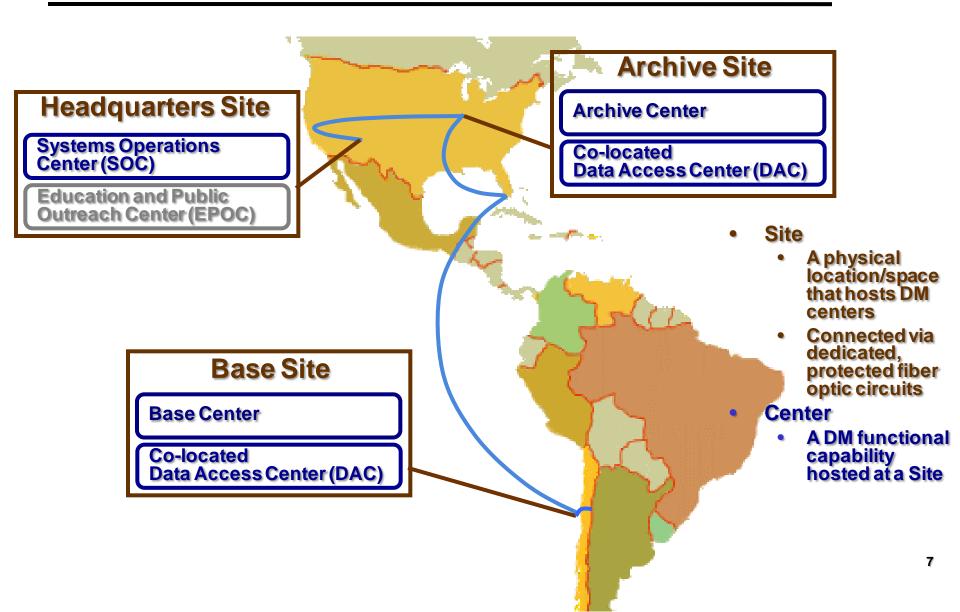




LSST will make tens of trillions photometric observations of tens of billions of objects

DM System is widely distributed

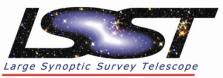




8

DM System relies on large-scale computational parallelism

Cluster Platform Pipeline With few exceptions, LSST Orchestration pipeline processing is Layer "embarassingly parallel" Pipeline Clipboard 3024 parallel ^{I/O} Stage image readouts O(10⁸) sky tiles Queue O(10⁹) objects Stage Data Trigger Event **Computational clusters** Queue Filesystem are well matched to the available parallelism Stage 5000 cores at Queue Event Base Broker ^{I/O} Stage 12000 (yr1) -33000 (yr10) cores at Archive Queue Inter-pipeline Stage **Middleware implements** Database Communication flexible Event Queue pipeline/production model ^{I/O} Stage of parallelism Data flow Control flow



DATA PRODUCTS

Application Layer -

Generates open, accessible data products with fully documented quality

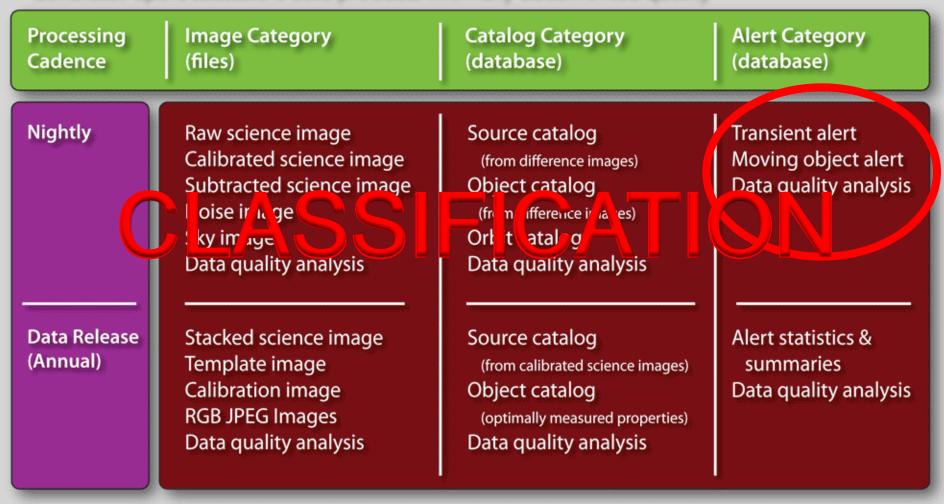
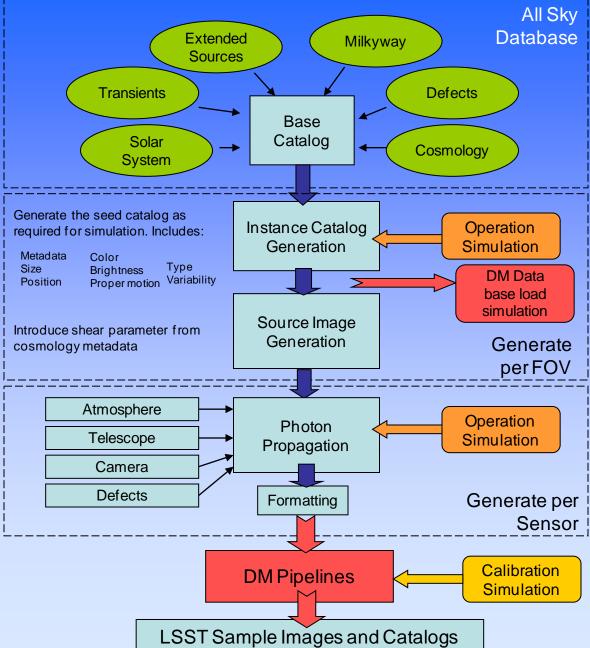
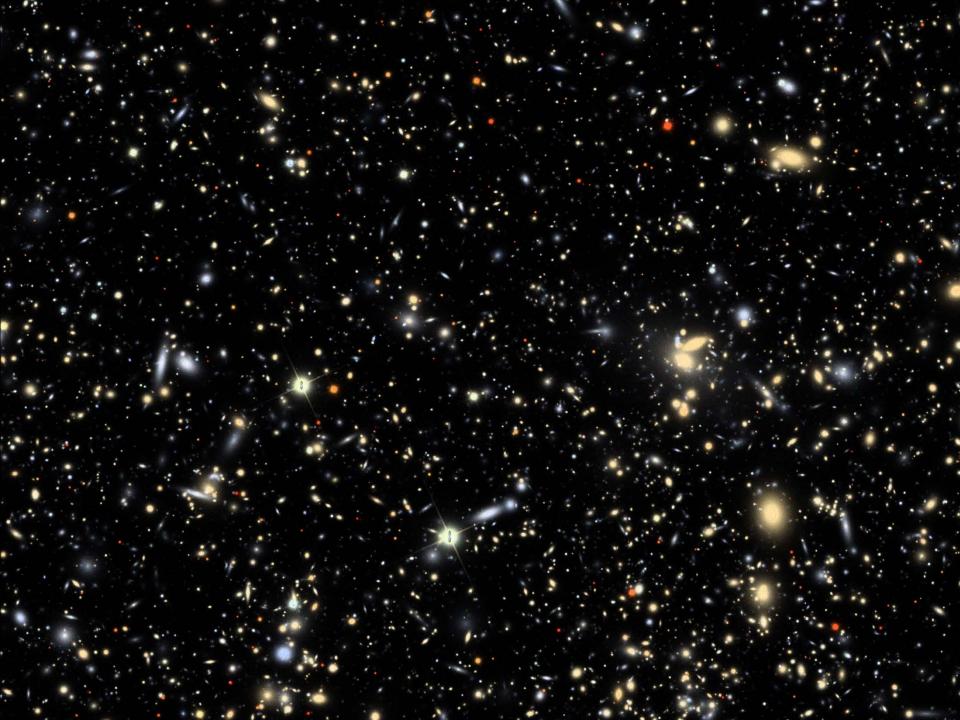


IMAGE SIMULATIONS



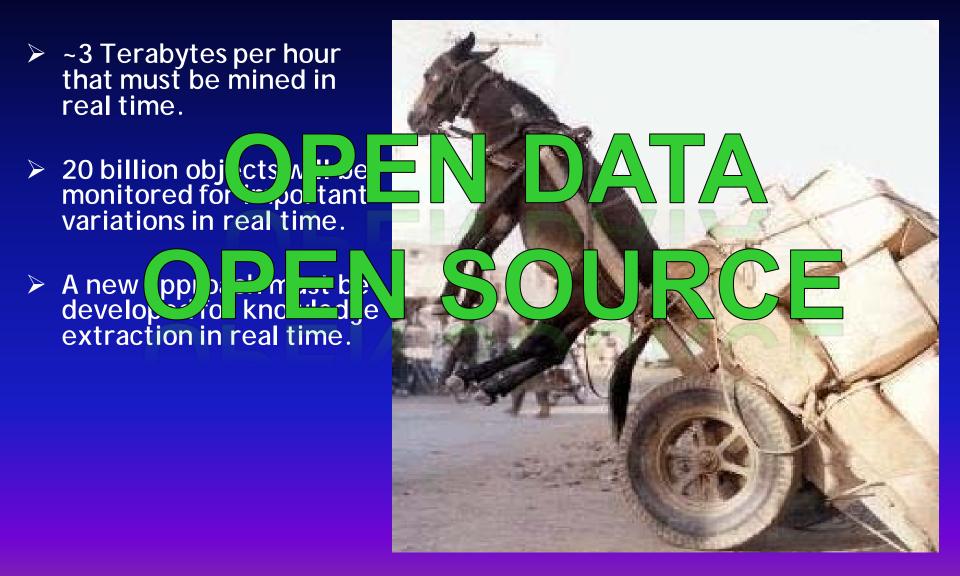


The Data Challenge

- 3 Terabytes per hour that must be mined in real time.
- 20 billion objects will be monitored for important variations in real time.
- A new approach must be developed for knowledge extraction in real time.

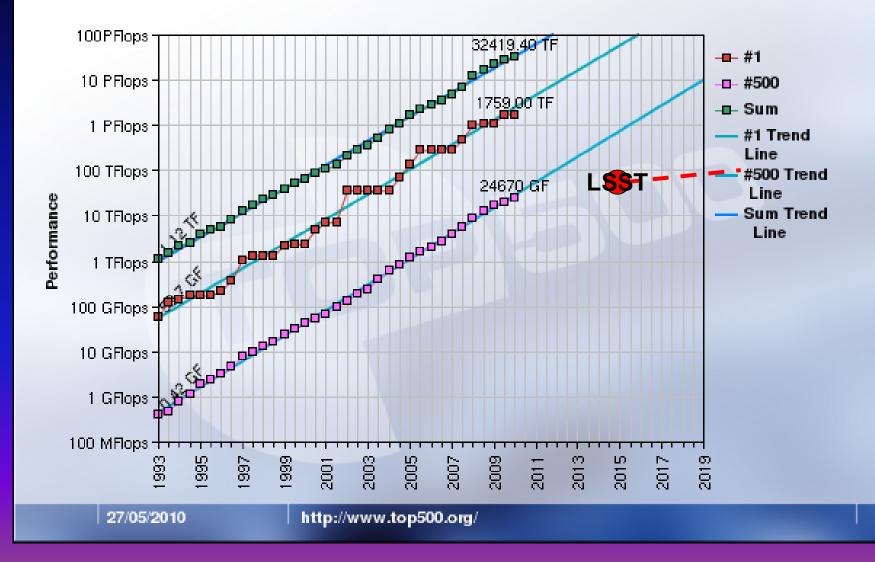


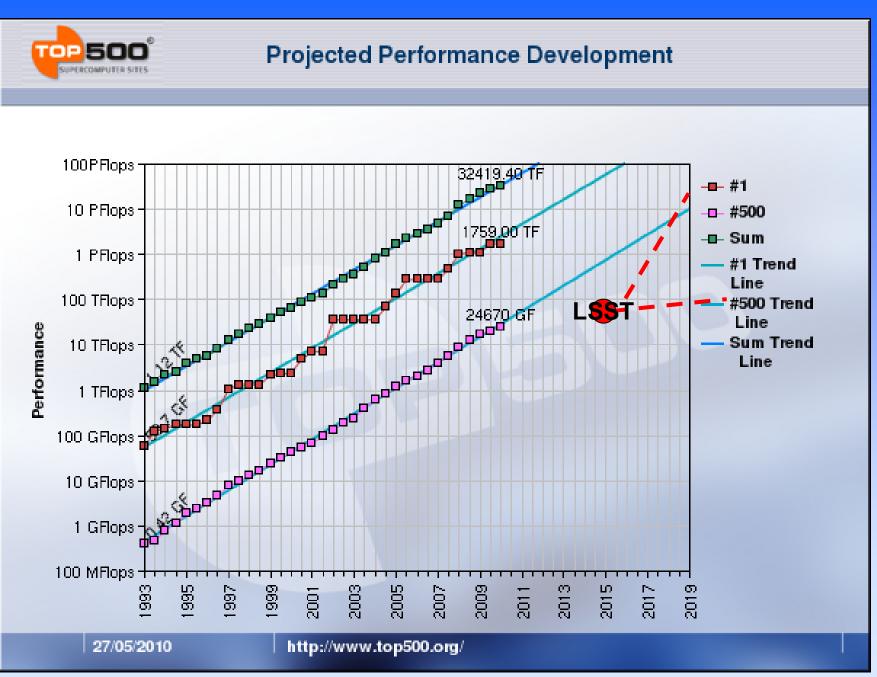
The Data Challenge





Projected Performance Development





Analytics

Complex computations

 100s of attributes per query

 Iterative, successively more restrictive
 Curiosity driven questions
 3 major query types

 Needle in haystack

- Correlations
- Time series

Science at the Limit

Much of the breakthrough science using surveys (imaging or spectroscopy) have occurred at the limits of the surveys

Sample incompleteness

Systematic errors

LSST Wide-Fast-Deep survey 4 billion galaxies with redshifts

Time domain:
1 million supervise
1 million galaxy lenses
1 billion moving objects
new phenomena

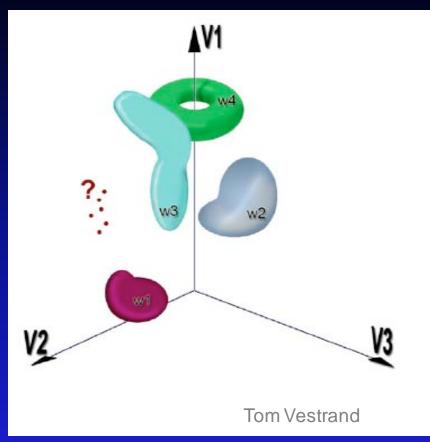
LSST Wide-Fast-Deep survey 4 billion galaxies with redshifts

Time domain:
1 million subcristed
1 million subcristed
1 million subcristed
1 billion moving objects
new phenomena

Major opportunity and challenge:

Discovering The Unexpected

- Characterize the known clustering
- Assign the new classification
- Discover the unknown (outlier detection)



Benefits of very large data sets:

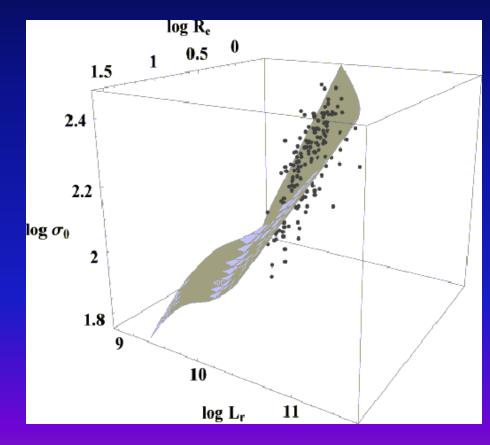
- best statistical analysis of "typical" events
- automated search for "rare" events

The dimension reduction problem:

Finding correlations and "fundamental planes" of parameters

The Curse of High Dimensionality !

- Are there combinations (linear or non-linear functions) of observational parameters that correlate strongly with one another?
- Are there eigenvectors or condensed representations (e.g., basis sets) that represent the full set of properties?



Automated discovery

Data exploration

DISCOVERING THE UNEXPECTED

This is required also for automated Data Quality Assessment

How To Learn More / Get Involved?

Open conference starting this year

1st public release

LSST lsst.org

Check out LSST database trac at
 <u>http://dev.lsstcorp.org/trac/wiki/LSSTDatabase</u>

> XLDB

- XLDB4 (Oct 6-7@SLAC)
- Read past XLDB reports
 <u>http://www-conf.slac.stanford.edu/xldb</u>
- Share your use cases, join the community

SciDB

- Check out http://scidb.org
- Try it out